EAS270. "The Atmosphere"	Quiz 1	21 Sept., 2009
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Professor: J.D. Wilson Time available: 15 mins Potential Value: 10%

**Instructions**: For all 10 questions, choose what you consider to be the best (or most logical) option, and use a pencil to mark that choice on the answer form. **Eqns/data given at back**. You may keep this quiz.

- 1. The troposphere, which extends from the base of the atmosphere to a height of about \_\_\_\_\_\_, is characterized by a steady \_\_\_\_\_\_ of climatological mean temperature with increasing height
  - (a) 80 kilometers; increase
  - (b) 50 kilometers; increase
  - (c) 10 kilometers; increase
  - (d) 50 kilometers; decrease
  - (e) 10 kilometers; decrease  $\checkmark \checkmark$
- 2. The two most abundant permanent gases in earth's homosphere are \_\_\_\_\_\_; and together they account for \_\_\_\_\_\_% of the air by volume
  - (a)  $N_2, O_2; 79\%$
  - (b) N<sub>2</sub>, O<sub>2</sub>; 99%  $\checkmark \checkmark$
  - (c)  $N_2$ , Ar ; 79%
  - (d)  $O_2$ , Ar ; 99%
  - (e)  $N_2, H_2O; 63\%$
- 3. Density of the atmosphere \_\_\_\_\_
  - (a) increases with increasing height
  - (b) does not change with height within the homosphere (lowest 80 km)
  - (c) decreases with increasing height  $\checkmark \checkmark$
  - (d) has the MKS unit kg  $m^{-1}$  ("kilograms per metre")
  - (e) has the MKS unit Pa ("Pascals")
- 4. About \_\_\_\_\_\_ % of atmospheric mass lies above the 700 hPa level
  - (a) 99
  - (b) 70 √√
  - (c) 50
  - (d) 30
  - (e) 1

- 5. The gas released to the atmosphere by volcanoes ("outgassing") is mostly composed of
  - (a) nitrogen and oxygen
  - (b) nitrogen and water vapour
  - (c) carbon dioxide and nitrogen
  - (d) water vapour and carbon dioxide  $\checkmark \checkmark$
  - (e) methane and ozone
- 6. Solar elevation above the horizon in Edmonton (latitude 53.5 degrees N), at solar noon on the day of the winter solstice, is \_\_\_\_\_ degrees
  - (a) 13 ✓✓
  - (b) 23.5
  - (c) 45
  - (d) 60
  - (e) 75
- 7. If earth's average surface temperature were to increase, the rate of emission of radiation energy from its surface would \_\_\_\_\_ and the wavelength  $(\lambda_{max})$  of the peak in the emission spectrum would shift towards \_\_\_\_\_ wavelengths
  - (a) increase, shorter  $\checkmark \checkmark$
  - (b) increase, longer
  - (c) decrease, longer
  - (d) decrease, shorter
  - (e) increase, redder

8. The numerical value of earth's "solar constant" is about \_\_\_\_\_

- (a) 9.8  $[m s^{-2}]$
- (b)  $0.5 \ [\mu m]$
- (c)  $4 \, [\mu m]$
- (d)  $500 \, [W \, m^2]$
- (e) 1370  $[{\rm W~m^{-2}}]$   $\checkmark\checkmark$

## For the remaining questions, please refer to the attached surface analysis.

- 9. The pressure change at the Alberta station farthest to the northwest (and reporting  $T = 6^{\circ}$ C,  $T_d = 3^{\circ}$ C) was \_\_\_\_\_
  - (a) a rise of 19 hPa
  - (b) a rise of 1.9 hPa  $\checkmark \checkmark$
  - (c) no change
  - (d) 74 hPa
  - (e) 7.4 hPa
- 10. Sea-level corrected pressure at the location marked by the large cross ( X, just north of the northern border of British Columbia) was about \_\_\_\_\_ hPa
  - (a) 10.14
  - (b) 1004
  - (c) 1014 ✓✓
  - (d) 1020
  - (e) 1027

## Equations and Data.

• 
$$p = \frac{M g}{A}$$

The pressure p [Pa] that results when a mass M [kg] of air overlies area A [m<sup>2</sup>], where  $g \approx 10 \text{ [m s}^{-2}$ ] is the acceleration due to gravity

• 1 hPa = 100 Pa

Pressure unit conversion. Sea-level pressure on earth is roughly 1000 hPa.

•  $L \uparrow = \epsilon \sigma T^4$ 

Stefan-Boltzmann law.  $L \uparrow [W m^{-2}]$ , the emitted longwave energy flux density (for which our textbook uses the symbol "I");  $\epsilon$ , the emissivity of the surface (dimensionless);  $\sigma = 5.67 \times 10^{-8}$  [W m<sup>-2</sup> K<sup>-4</sup>], the Stefan-Boltzmann constant; T [K], the surface temperature.

•  $\lambda_{max} = \frac{2900}{T}$ 

Wien's displacement law.  $\lambda_{max}$  [µm], the wavelength at which the peak in the emission spectrum occurs; T [K], the temperature of the emitting surface.

•  $\theta = 90 - \Phi_{lat} + \phi_{sol.dec}$ 

The solar elevation  $\theta$  at solar noon, at a location with latitude  $\Phi_{lat}$ , at the time of year when solar declination is  $\phi_{sol.dec}$ . Latitude is negative in the southern hemisphere; and solar declination is negative during northern hemisphere winter.



Figure 1: CMC surface analysis, 12Z Sept. 10, 2009